

CFD Modeling of NASA's ARES Platform

Joshua Coltrane

**Arkansas-Oklahoma Center for Space
and Planetary Science**

July 29th, 2004

University of Arkansas



Computational Fluid Dynamics

- * Cheaper and less time consuming than building models and testing in wind-tunnels.
- * Processing speeds are allowing for more intense computations
- * More flexibility in testing, e.g. Mars



FELISA Grid Generation for ARES

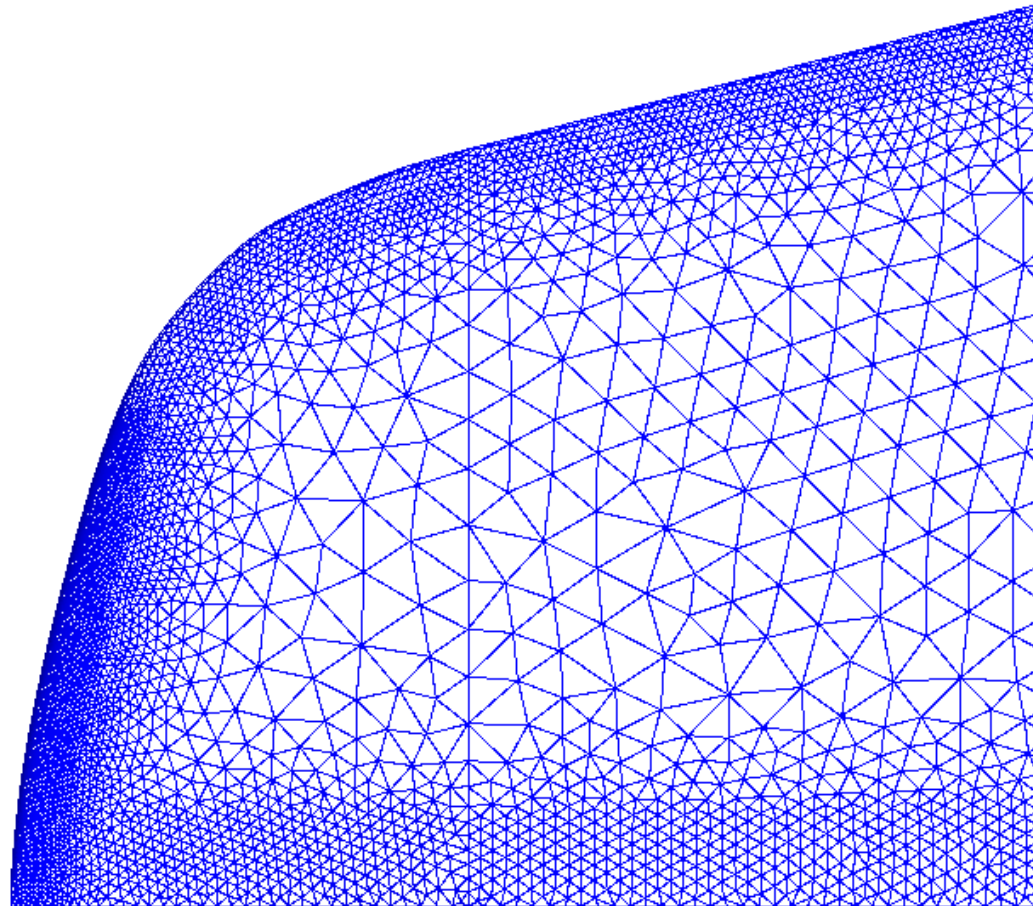
Files needed:

.con

.bco

.bac

.sur



The Solver

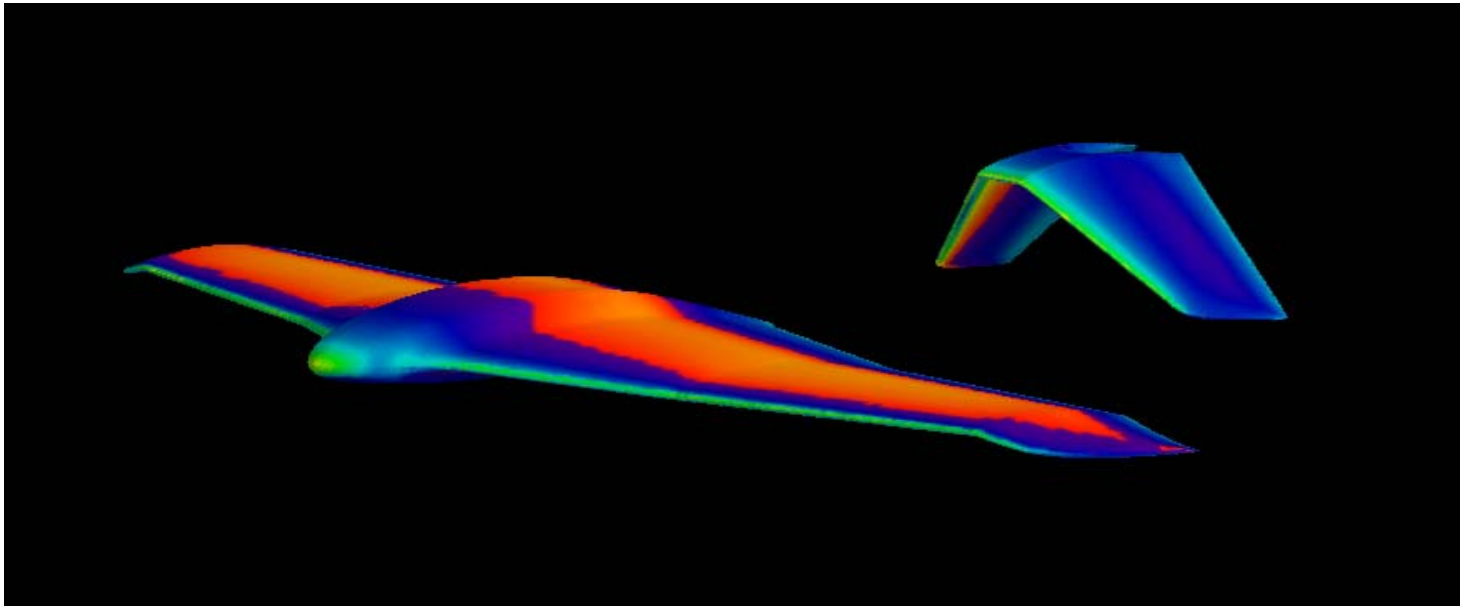
Euler3d is an modified fluid flow solver designed towards handling non-inertial problems.

The output is in the form of a loads file containing the iteration number, time step, forces and moments.

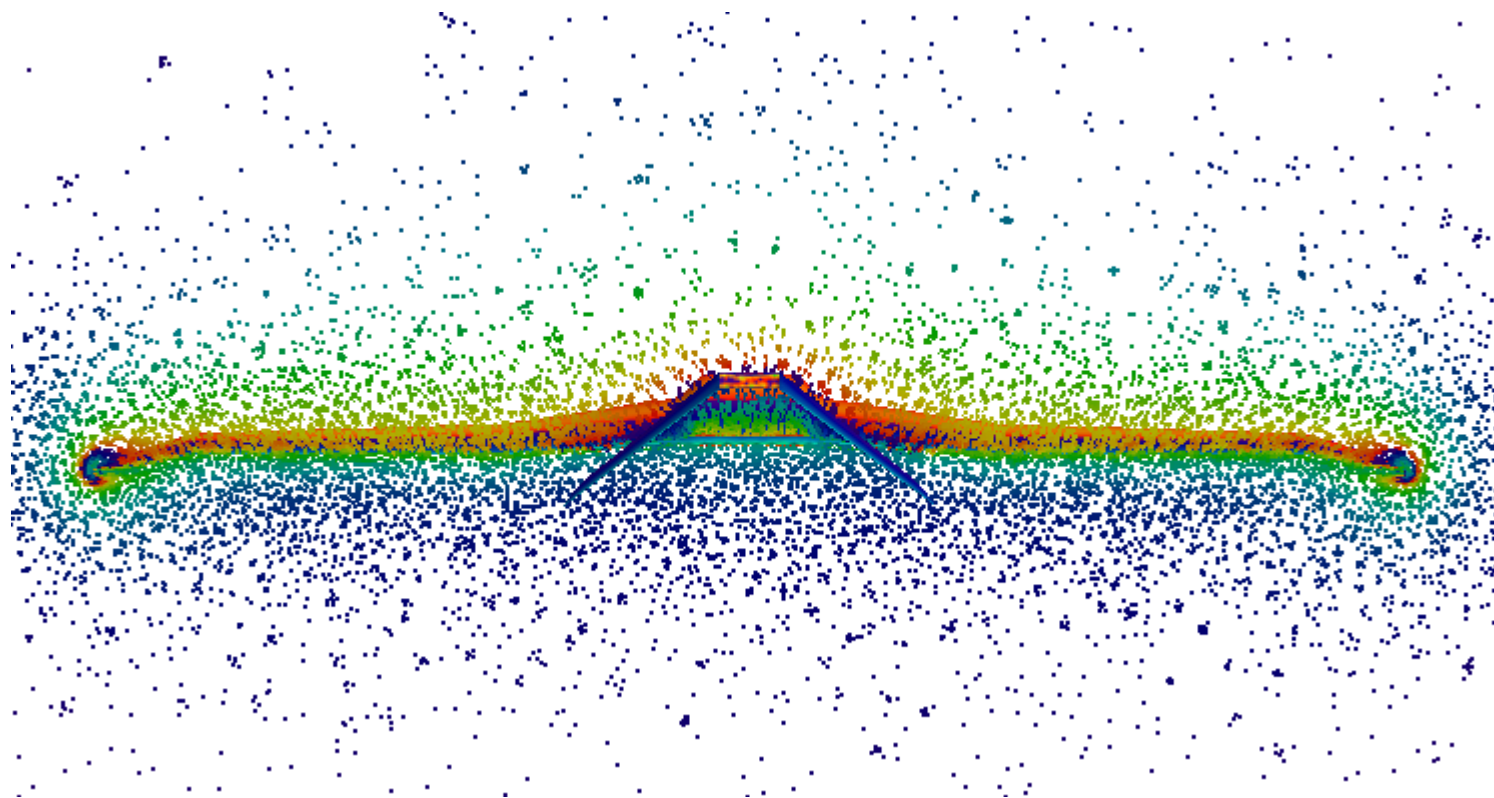
The loads file does iterations until it converges on a set of forces and moments.

4996	5.00E+02	8.80E-02	-3.64E-03	2.20E+00	-7.33E-03	-2.93E+00	-1.48E-02
4997	5.00E+02	8.80E-02	-3.64E-03	2.20E+00	-7.33E-03	-2.93E+00	-1.48E-02
4998	5.00E+02	8.81E-02	-3.64E-03	2.20E+00	-7.33E-03	-2.93E+00	-1.48E-02
4999	5.00E+02	8.82E-02	-3.64E-03	2.20E+00	-7.32E-03	-2.93E+00	-1.48E-02
5000	5.00E+02	8.82E-02	-3.64E-03	2.20E+00	-7.32E-03	-2.93E+00	-1.48E-02

Now that we have these numbers we can view the flow governed by the parameters set in the controls file over ARES in the glplot3d program...



A velocity vectors cut showing vortices behind the wingtips...



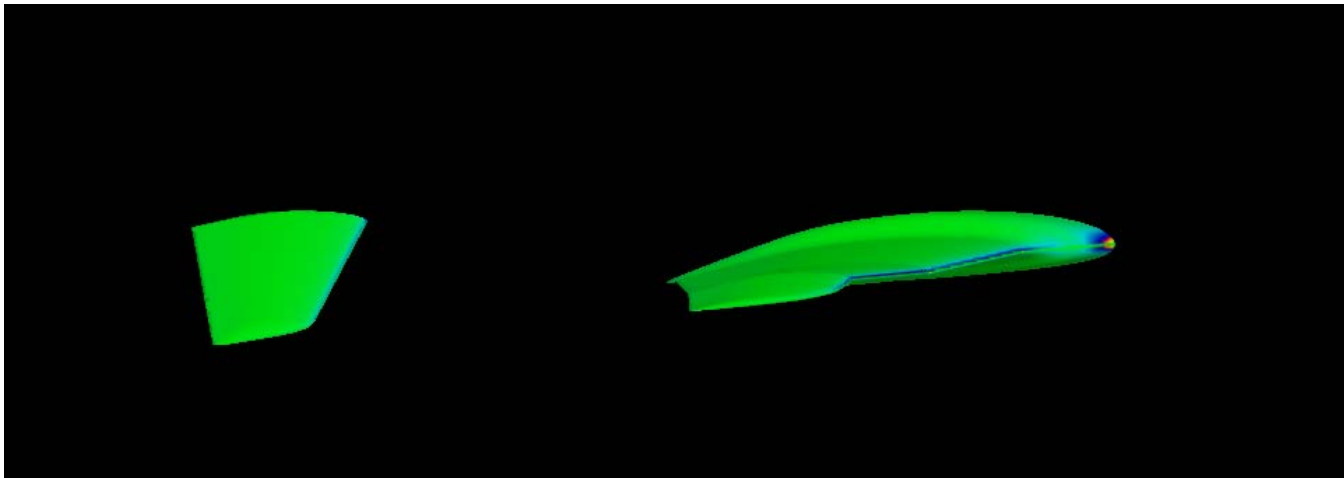
The output can also be used to find the C_L , C_D or the pitching moment around the center of gravity at different angles of attack.

$$C_L = \frac{\text{Lift}}{(\frac{1}{2}\rho v^2 S)}$$

$$C_D = \frac{\text{Drag}}{(\frac{1}{2}\rho v^2 S)}$$

$$M = \frac{1}{2}\rho v^2 S c C_M$$

$$\text{Dynamic Pressure} = \frac{1}{2}\rho v^2$$

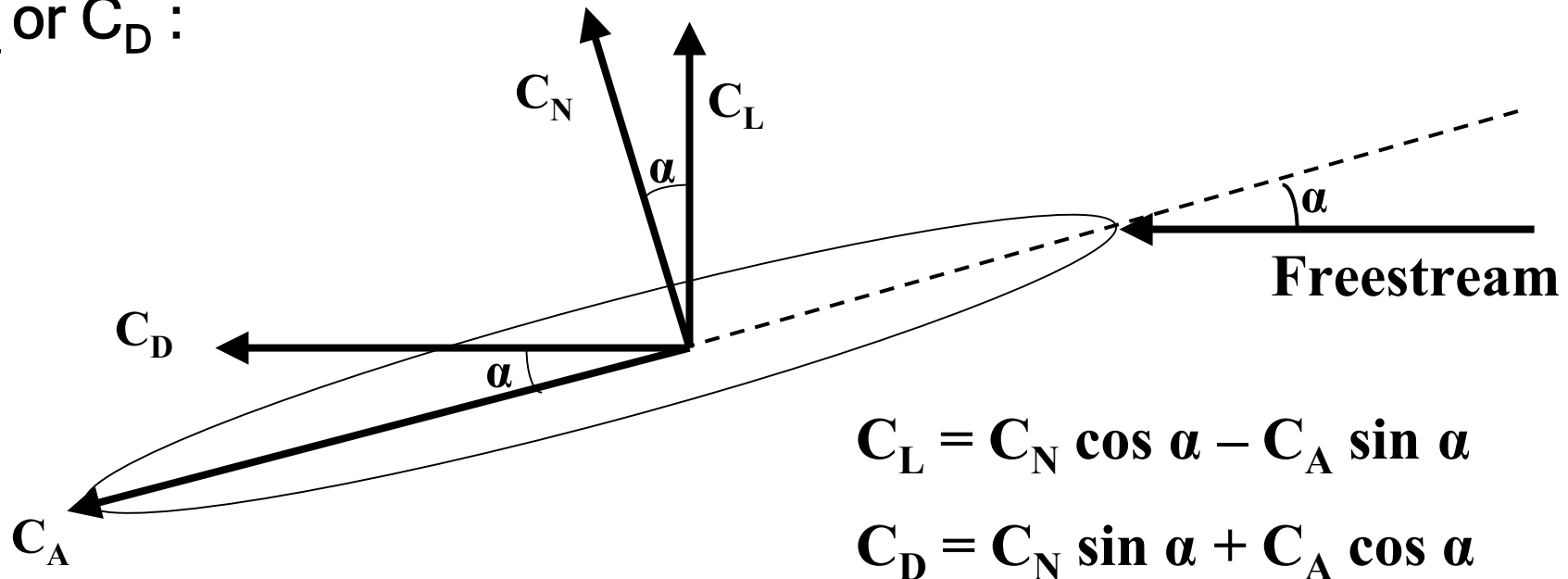


To arrive at the C_L or C_D , the force coefficients are taken from the last iteration in the loads file:

5000	5.00E+02	<u>8.82E-02</u>	<u>-3.64E-03</u>	2.20E+00	-7.32E-03	-2.93E+00	-1.48E-02
------	----------	-----------------	------------------	----------	-----------	-----------	-----------

Planform Area (6.55 m²)

The forces, axial in the x direction and normal in the z direction, then go through some trigonometry to become C_L or C_D :



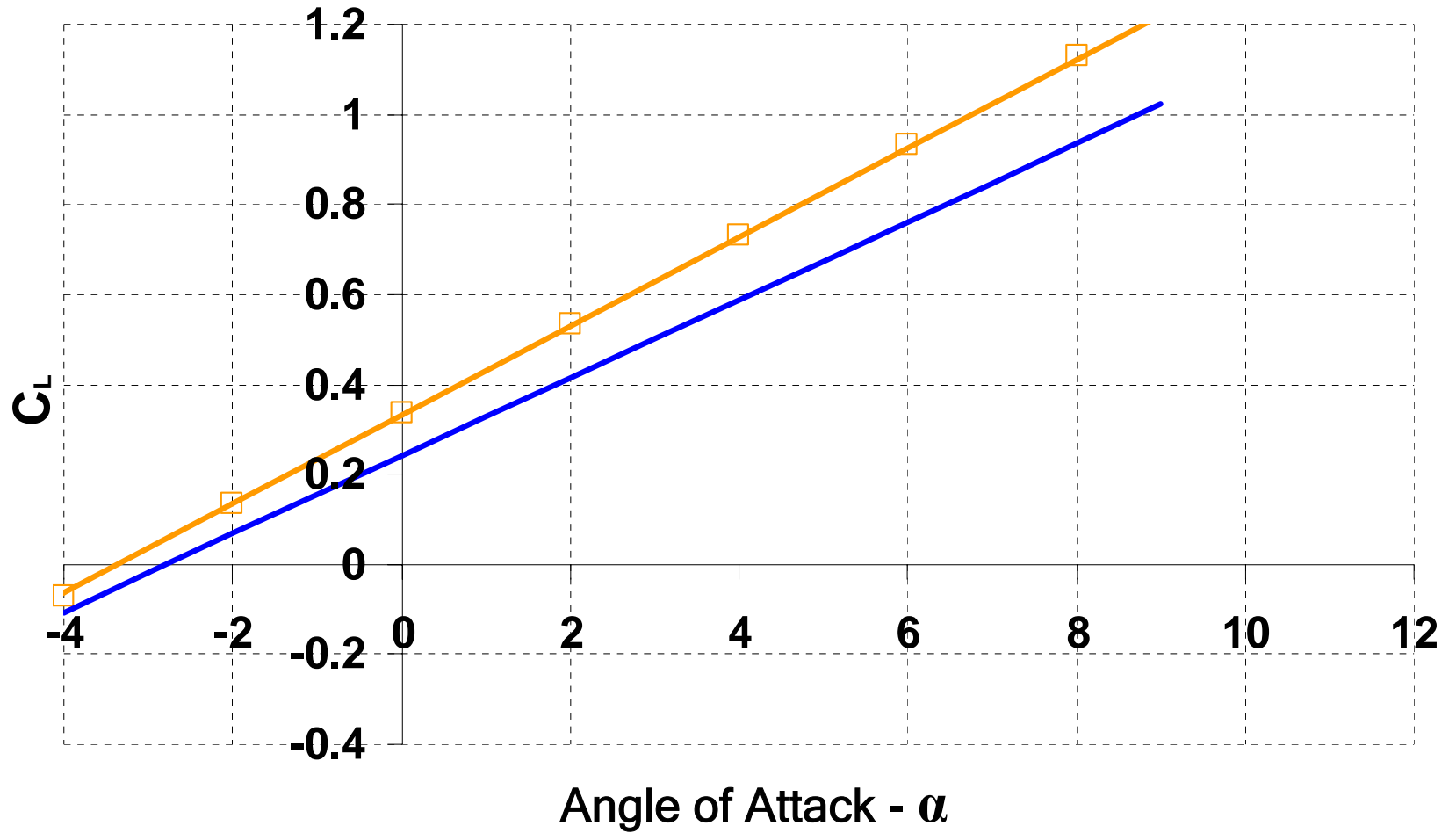
$$C_L = C_N \cos \alpha - C_A \sin \alpha$$

$$C_D = C_N \sin \alpha + C_A \cos \alpha$$

C_L Data for Modeled ARES

Euler3d

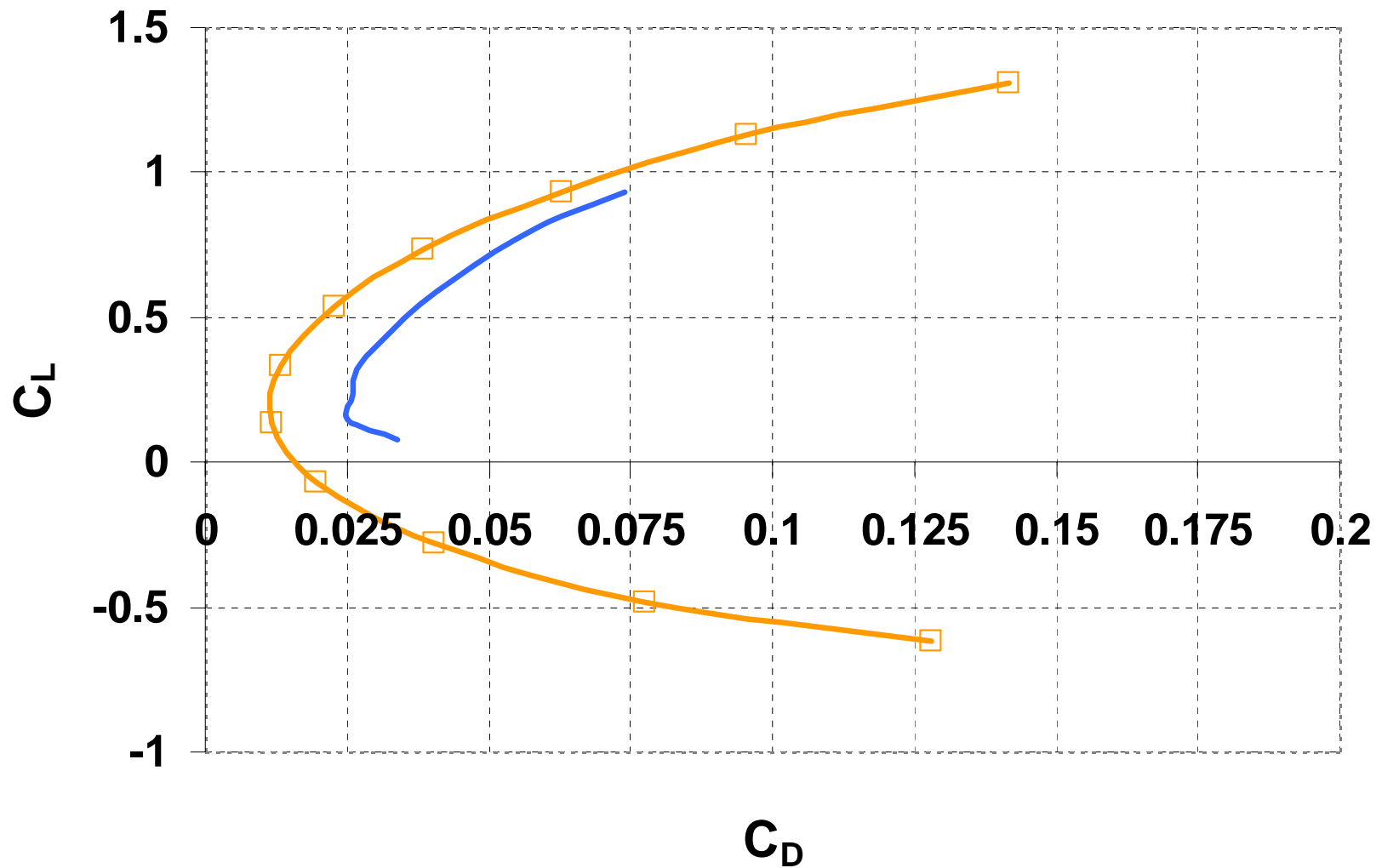
Experimental



C_L & C_D Data for Modeled ARES

Euler3d

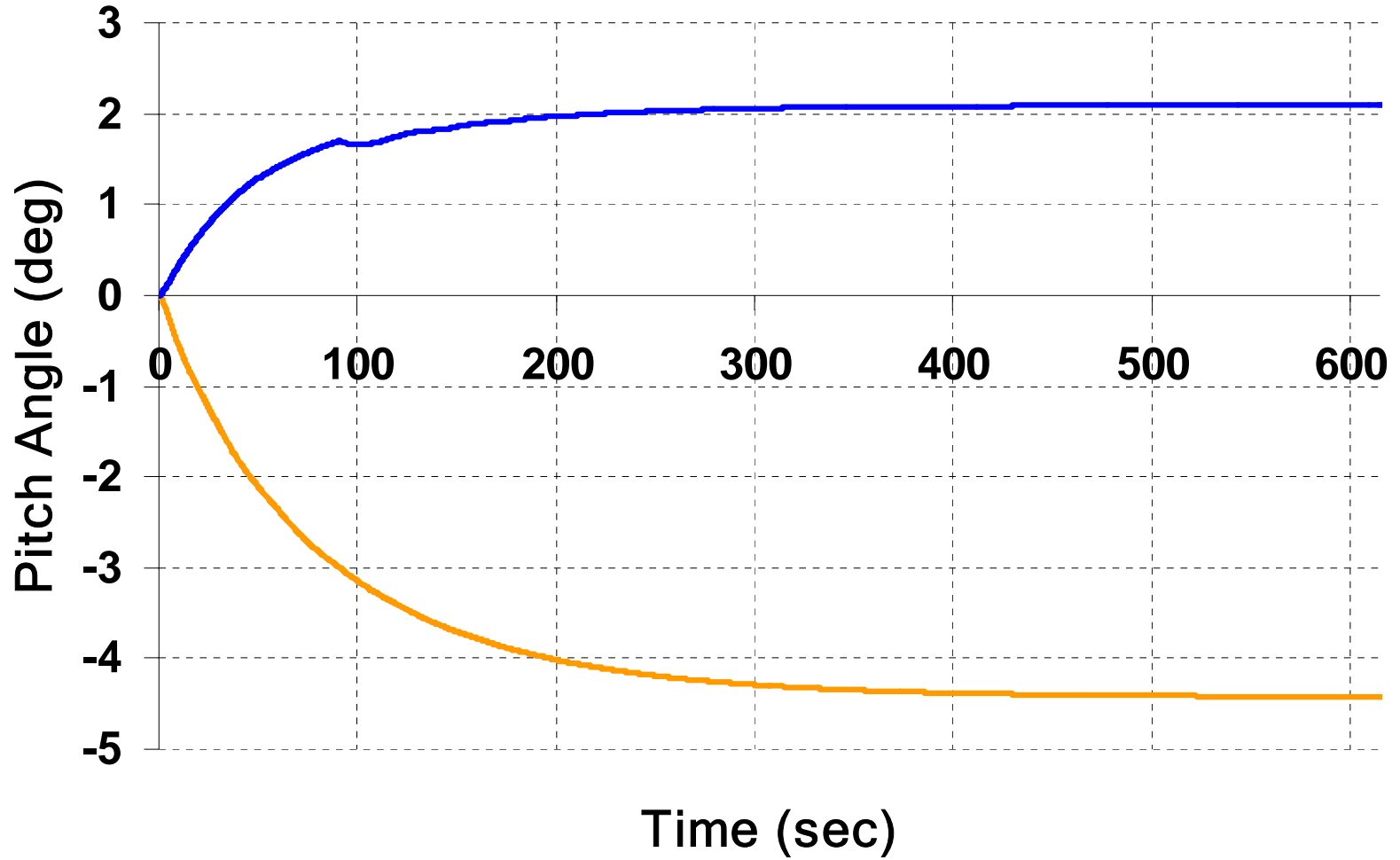
Experimental



Angle Convergence

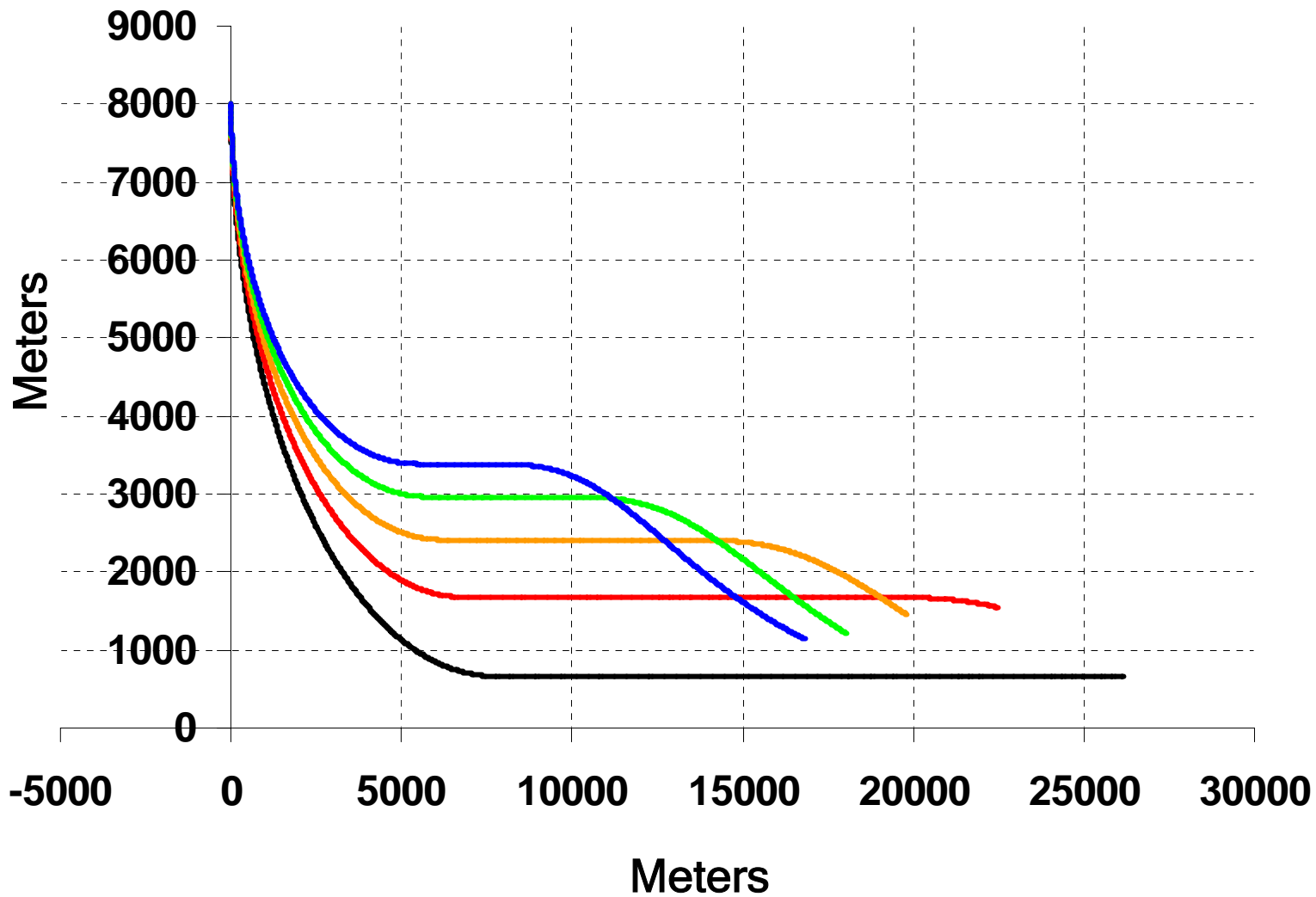
0°

5°



Pullout Trajectories

12° 14° 16° 18° 20°



The pitching moment is found by dividing the moment coefficient in the y direction by the planform area and mean aerodynamic chord.

5000	5.00E+02	8.82E-02	-3.64E-03	2.20E+00	-7.32E-03	<u>-2.93E+00</u>	-1.48E-02
------	----------	----------	-----------	----------	-----------	------------------	-----------

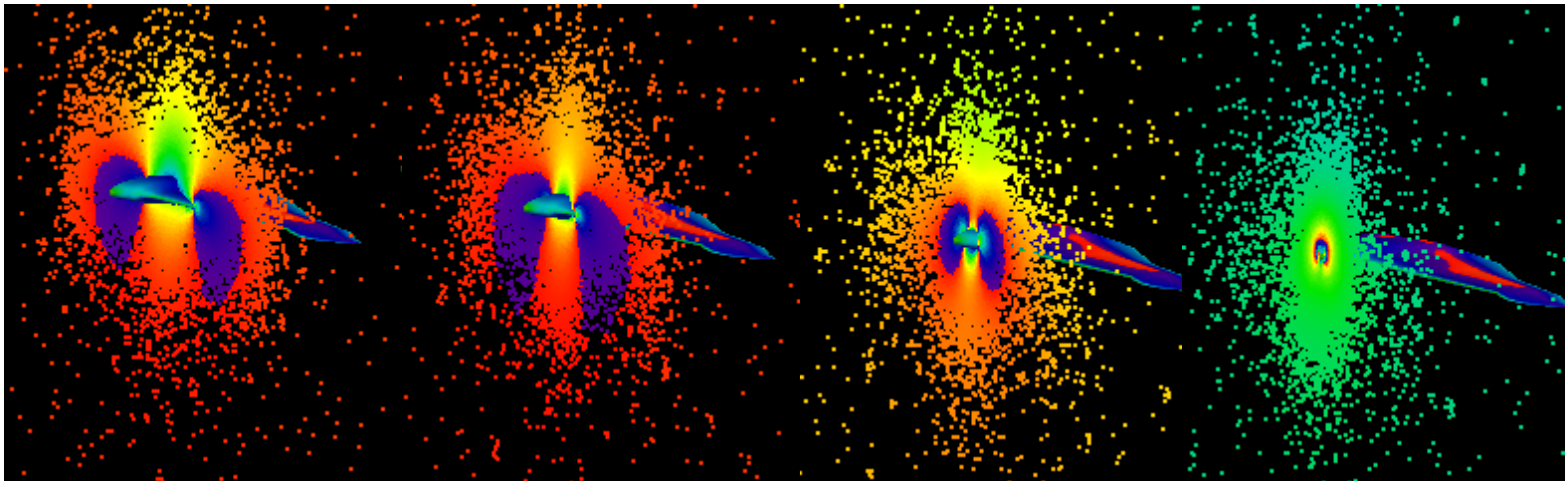
6.55 m² * 1.05 m

$$M = \frac{1}{2} \rho v^2 S c C_M$$

This would give a pitching moment coefficient of -.426

Summary

- * Cheaper and less time consuming than building models and testing in wind-tunnels.
- * Processing speeds are allowing for more intense computations
- * More flexibility in testing, e.g. Mars



CFD Modeling of NASA's ARES Platform

- **Joshua Coltrane**
- **Arkansas-Oklahoma Center for Space and Planetary Science**
- **July 29th, 2004**
- **University of Arkansas**

